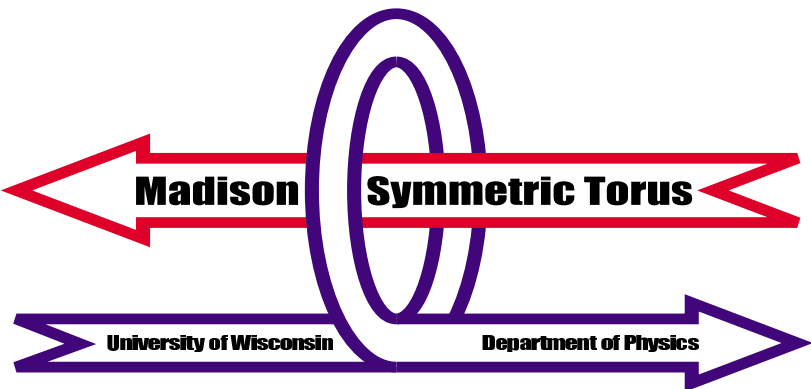


# Momentum transport from tearing modes

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# Motivations

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- Both in laboratory and astrophysical plasmas, there are situations in which momentum transport and flow phenomena can not be explained by classical processes.
- Magnetic fluctuations arising from MHD instabilities are believed to have an important role in transporting momentum.
- In astrophysical plasmas, momentum is believed to be transported due to flow-driven instabilities (MRI).
- Here, we investigate whether current-driven reconnection can transport momentum.
- Is there correlation between the anomalous momentum transport and the tearing fluctuations?
- Whether spontaneous and/or driven reconnection can cause momentum transport?

# Outline

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- Momentum transport from a [single tearing mode](#)
  - Analytical calculations of quasilinear  $\langle \tilde{J} \times \tilde{B} \rangle$  torque in the presence of mean flow
  - Computation of  $\langle \tilde{J} \times \tilde{B} \rangle$  torque
- Momentum transport from [multiple tearing modes](#) – role of coupled spontaneous and driven reconnection

# Summary

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## Single tearing mode

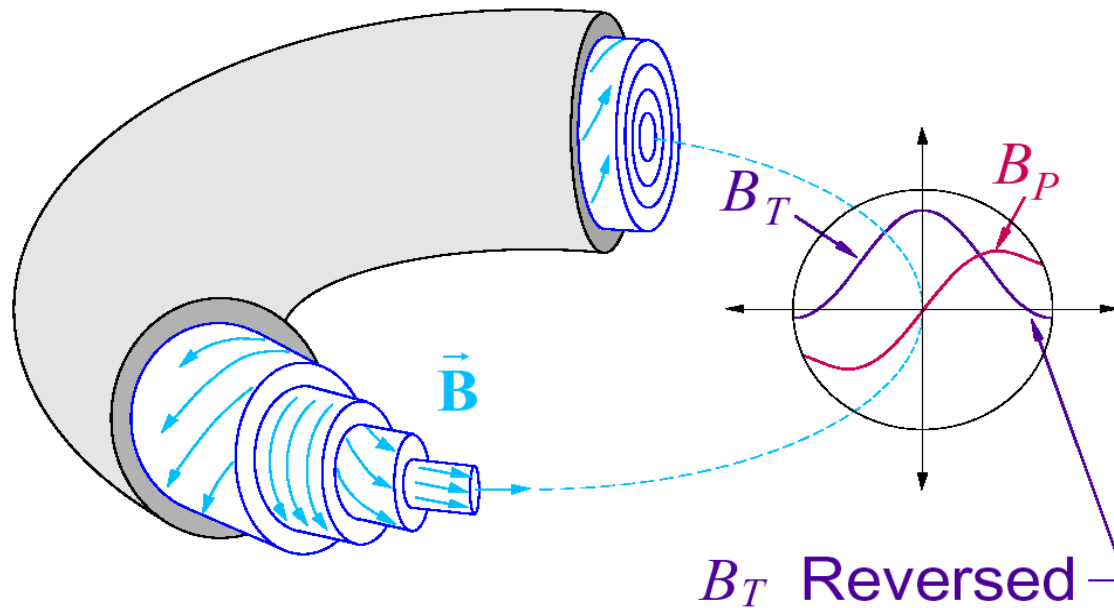
- In the absence of mean flow
  - $\langle \tilde{J} \times \tilde{B} \rangle = 0$
- In the presence of mean flow
  - In the outer region,  $\langle \tilde{J} \times \tilde{B} \rangle \approx 0$
  - In the inner layer,  $\langle \tilde{J} \times \tilde{B} \rangle \neq 0$

## Multiple tearing modes

- $\langle \tilde{J} \times \tilde{B} \rangle \neq 0$   
nonlinear mode coupling enhances momentum transport.

# Plasma flow is affected by the MHD turbulence in lab plasmas.

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- **Tearing fluctuations** (about a few % of mean field) play an important role in sustaining the RFP configuration (e.g. current transport, sawteeth oscillations).

## Experimental observations

- There is large mean flow (about a few percent of  $V_A$ ) in RFP.
- The flow decelerations in RFP are coherent with the relaxation events (sawteeth oscillations).

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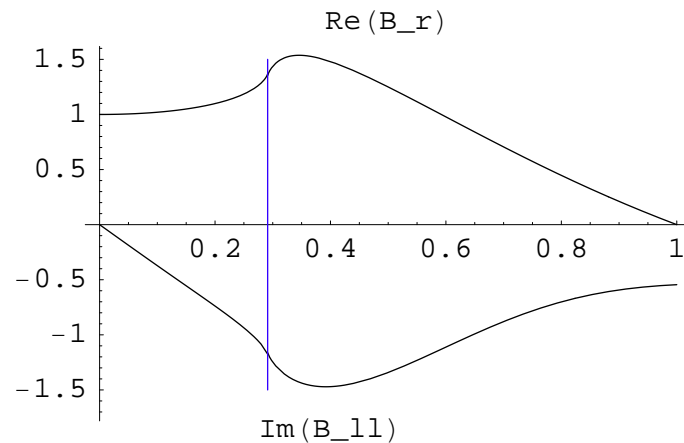
# Analytical calculations of quasilinear torque

# The outer solutions are modified by the mean flow profile.

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- In the absence of mean flow

- The linearized outer solutions are obtained from Newcomb equation.



- $(\tilde{\mathbf{J}} \times \tilde{\mathbf{B}})_{||} = \frac{1}{r} \frac{\partial}{\partial r} \langle r \tilde{B}_{||} \tilde{B}_r \rangle = 0$

- $\tilde{B}_{||}$  and  $\tilde{B}_r$  are out of phase.

- In the presence of mean flow

- We solve linearized outer equations with mean flow;

$$Q(r, \theta, z, t) = Q(r) \exp(i \omega t + i (m \theta + k z))$$

$$\mathbf{B} = B_\theta(r) \mathbf{e}_\theta + B_z(r) \mathbf{e}_z$$

$$\mathbf{V}_0 = V_z(r) \mathbf{e}_z$$

$$(r \tilde{V}_r)'' + c1(r \tilde{V}_r)' + c2(r \tilde{V}_r) = 0$$

# The outer solutions are modified by the mean flow profile.

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$$c1 = \frac{1}{g} \frac{dg}{dr} - \frac{kV'_z}{\Omega}$$
$$c2 = \frac{1}{g} \left[ \frac{d}{dr} \left( \frac{kV'_z g}{\Omega} \right) + \frac{2k^2 B_\theta V'_z G}{\Omega^2 H} - \frac{d}{dr} \left( \frac{2mB_\theta F}{r\Omega H} \right) \right. \\ \left. + \frac{2B_\theta}{r\Omega} \left( \frac{B_\theta}{r} \right)' + \frac{4k^2 B_\theta^2}{r\Omega(1-M^2)H} - \frac{F^2(1-M^2)}{r\Omega} - \frac{2B_\theta^2 kV'_z}{r^2 \Omega^2} \right]$$

$$g = \frac{rF^2(1-M^2)}{\Omega H}, \quad H = m^2 + k^2 r^2$$
$$F = \frac{mB_\theta}{r} + kB_z, \quad G = -\frac{mB_z}{r} + kB_\theta$$

Earlier studies by Gatto, Terry and Hegna (2002)

# The single mode Lorentz force is small in the outer region.

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$$\tilde{B}_{||} = i(1 - M^2)\tilde{B}_{||}^{(0)}$$

$$\tilde{B}_{||}^{(0)} = \left[ \frac{Fr(r\tilde{B}_r)' - rG\lambda(r\tilde{B}_r)}{B(m^2 + k_z^2 r^2)} \right] = \text{solution without flow}$$

$$M = \sqrt{\rho} \frac{(\omega + kV_z)}{K \cdot B} \rightarrow \text{Machnumber}$$

$$\omega = \omega_r + i\gamma, \quad \omega_r = -kV_z|_{(r=r_s)}$$

The single mode Lorentz term is small due to the small growth rate in the outer region.

$$\boxed{(\tilde{J} \times \tilde{B})_{||} = \frac{4\gamma}{r} \left[ \rho \frac{(\omega_r + kV_z)}{(K \cdot B)^2} r \tilde{B}_{||}^{(0)} \tilde{B}_r \right]'}$$

# The inner layer equations are solved in cylinder with mean flow

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The set of linearized MHD momentum, adiabatic pressure, and induction equations in the presence of a mean flow velocity are;

$$\begin{aligned}\rho\omega^2\xi + \rho\omega(\mathbf{v}_0 \cdot \nabla)\xi + \rho\omega(\xi \cdot \nabla)\mathbf{v}_0 &= (\nabla \times \tilde{\mathbf{B}}) \times \mathbf{B} + \mathbf{J} \times \tilde{\mathbf{B}} - \nabla\tilde{p} \\ \tilde{p} &= -(\xi \cdot \nabla p) - \gamma p \nabla \cdot \xi \\ \tilde{\mathbf{B}} - \eta/\omega \nabla^2 \tilde{\mathbf{B}} &= \nabla \times (\xi \times \mathbf{B}) + 1/\omega \nabla \times (\mathbf{v}_0 \times \tilde{\mathbf{B}})\end{aligned}$$

1- adopt CGJ tearing ordering ( $\omega \propto \eta^{3/5}$ )

2- consider small flow shear

•  $\omega \rightarrow \epsilon^3$ ,  $\eta \rightarrow \epsilon^5$ ,  $(r - r_s) \rightarrow \epsilon^2 x$ ,  $v_0 \approx \eta^{1/5}$  (the inertial regime)

•  $g = (k \cdot \mathbf{v}_0)'_{(r_s)} x \rightarrow \epsilon^3$ ,

$\boxed{(k \cdot \mathbf{v}_0)'_{(r_s)}} \rightarrow$  flow shear

## The inner layer equations are solved in cylinder with mean flow

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The inner resistive layer equations in the presence of mean flow :

$$\Psi_2'' = \Omega[X\Xi + (1 + i\frac{G'X}{\Omega})\Psi_0]$$

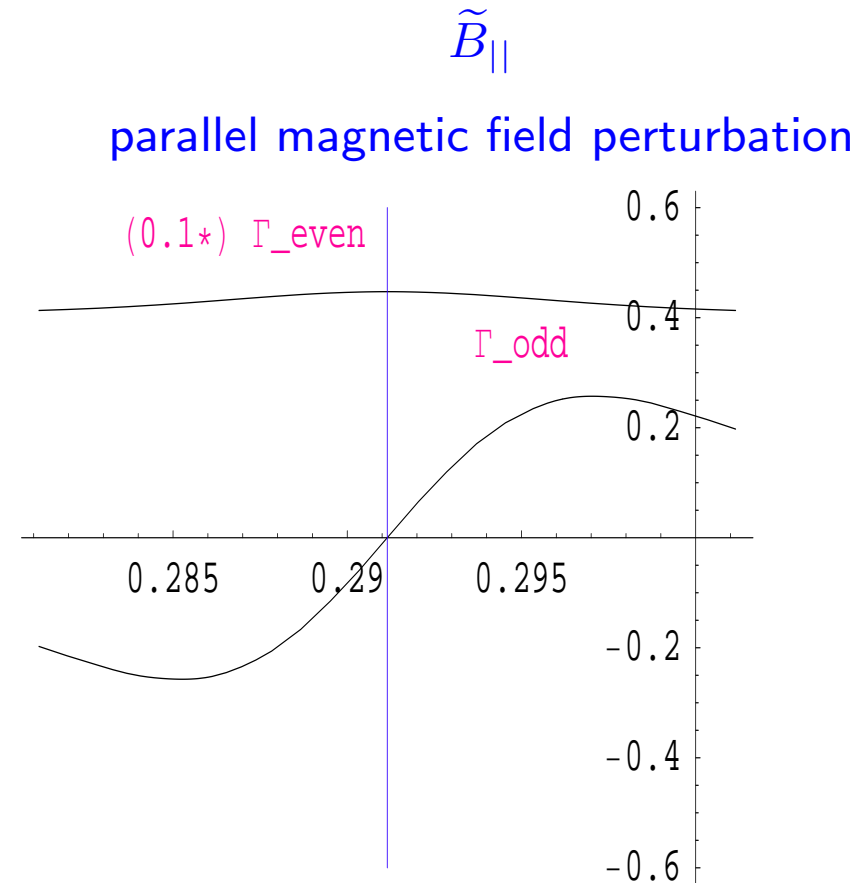
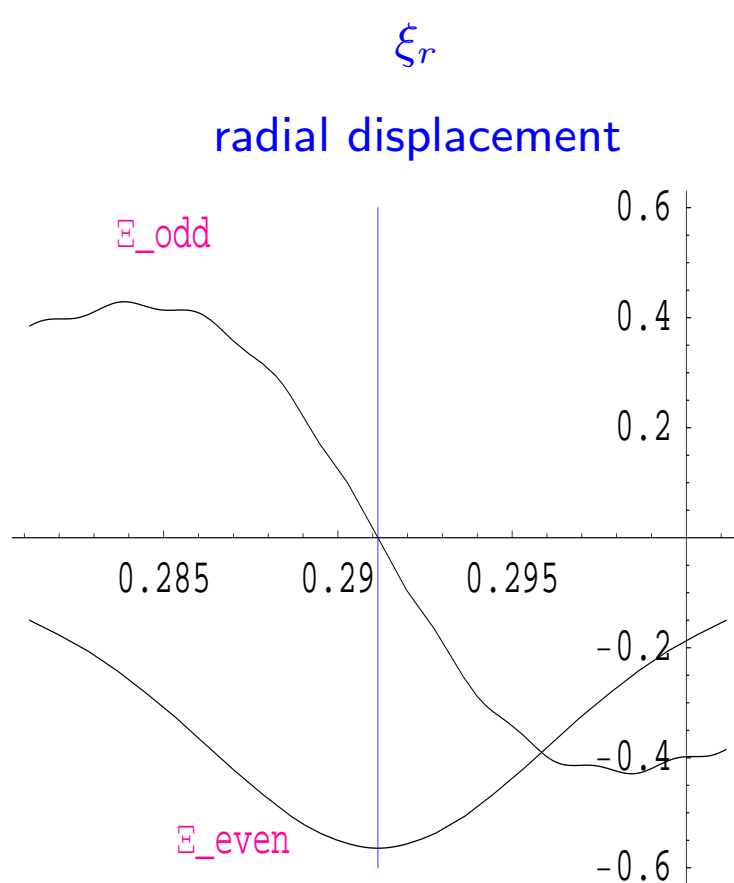
$$(1 + i\frac{G'X}{\Omega})\Xi'' = \frac{X^2\Xi}{\Omega} - \frac{\Upsilon}{\Omega^2} + [\frac{X}{\Omega}(1 + i\frac{G'X}{\Omega}) - \frac{J_p}{\Omega^2}]\Psi_0$$

$$(1 + i\frac{G'Z}{\Omega^{3/4}})\Upsilon'' - \left(Z^2 + \frac{2l^6}{\gamma\beta\Omega^{-3/2}}(1 + i\frac{G'Z}{\Omega^{3/4}})\right)\Upsilon - \left(\frac{S}{\Omega^{-3/2}}(1 + i\frac{G'Z}{\Omega^{3/4}}) + i\frac{G'G_pZ}{\Omega^{-3/4}}\right)\Xi + \left(JZ^2 + \frac{2Jl^6}{\gamma\beta\Omega^{-3/2}}(1 + i\frac{G'Z}{\Omega^{3/4}}) - i\frac{G'G_p}{\Omega^{-1/2}}(1 + i\frac{G'Z}{\Omega^{3/4}})\right)\Psi_0 = 0$$

We solve for  $\Xi^0$ ,  $\Xi^1$ ,  $\Upsilon^0$  and  $\Upsilon^1$  to calculate quasilinear  $\langle \tilde{J} \times \tilde{B} \rangle$  torque.

The inner layer zeroth order solutions are used to calculate the Lorentz term.

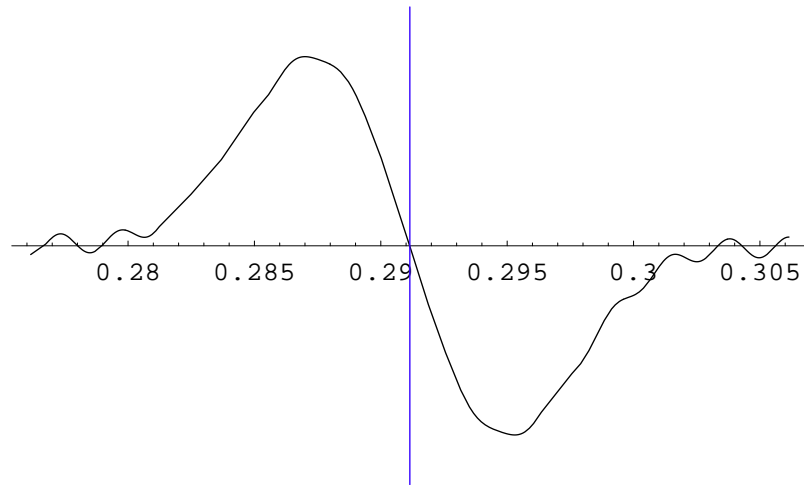
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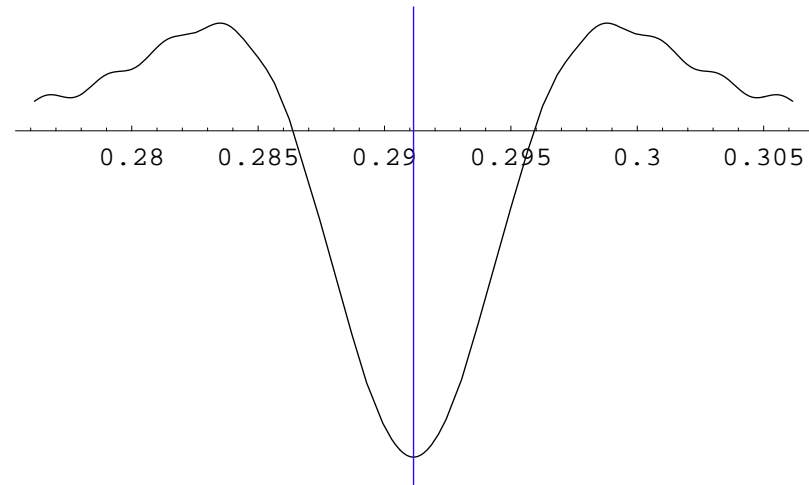
# The analytical quasilinear $\langle \tilde{J} \times \tilde{B} \rangle$ term in the presence of flow shear.

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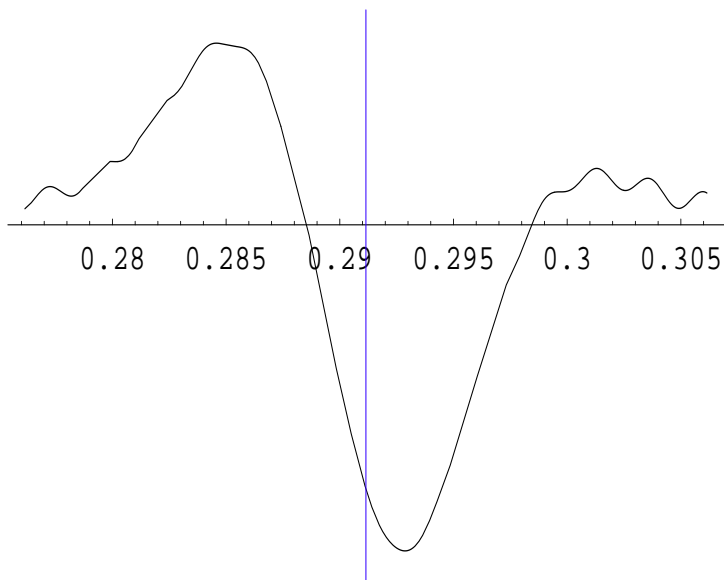
odd  $\langle \tilde{J} \times \tilde{B} \rangle$



even  $\langle \tilde{J} \times \tilde{B} \rangle$



total  $(J \times B)_{||}$  w/ flow



$$\langle \text{====} \text{ total } \langle \tilde{J} \times \tilde{B} \rangle$$

$$T = \int r \langle \tilde{J} \times \tilde{B} \rangle_z$$

$$dr = 0$$

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# Single mode computations

## 3-D resistive MHD code

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- The 3-D resistive nonlinear MHD code DEBS is used.

$$\begin{aligned}\frac{\partial \bar{A}}{\partial t} &= S \bar{V} \times \bar{B} - \eta \bar{J} \\ \rho \frac{\partial \bar{V}}{\partial t} &= -S \rho \bar{V} \cdot \nabla \bar{V} + S \bar{J} \times \bar{B} + \nu \nabla^2 \bar{V} + \mathbf{F}(r) \\ \bar{B} &= \nabla \times \bar{A} \\ \bar{J} &= \nabla \times \bar{B}\end{aligned}$$

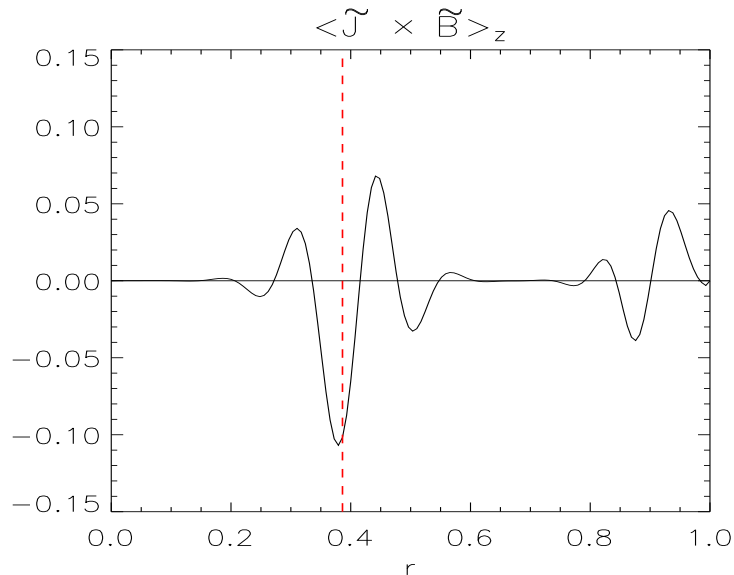
where  $S = \frac{\tau_R}{\tau_A}$  ( Lundquist number ) ,  $\nu = \frac{\tau_R}{\tau_{vis}}$  ( magnetic Prandtl number = viscosity / resistivity ).

- With an axial ad-hoc momentum source term  $\mathbf{F}(r)$  added to the momentum equation.

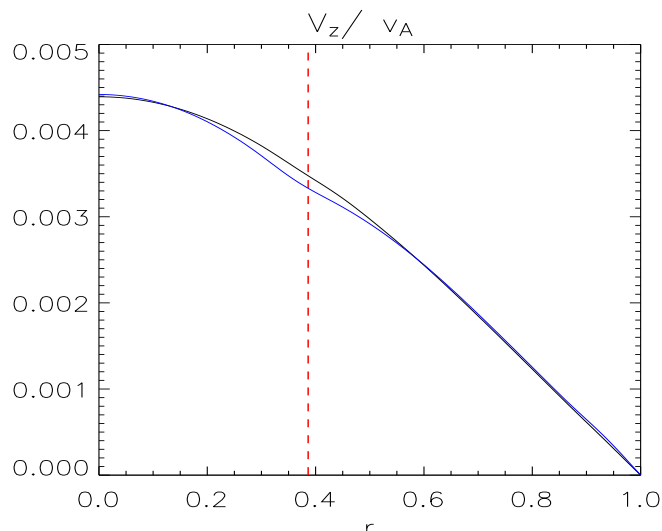
# Single tearing mode does transport momentum

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during the linear growth



- Single helicity  $\langle \tilde{J} \times \tilde{B} \rangle$  from computation agrees with the analytical calculations.



- Quasi-linear Lorentz force from the single mode perturbations affect the flow profile.

$$\langle \tilde{V}_z \rangle = V_z/V_A$$

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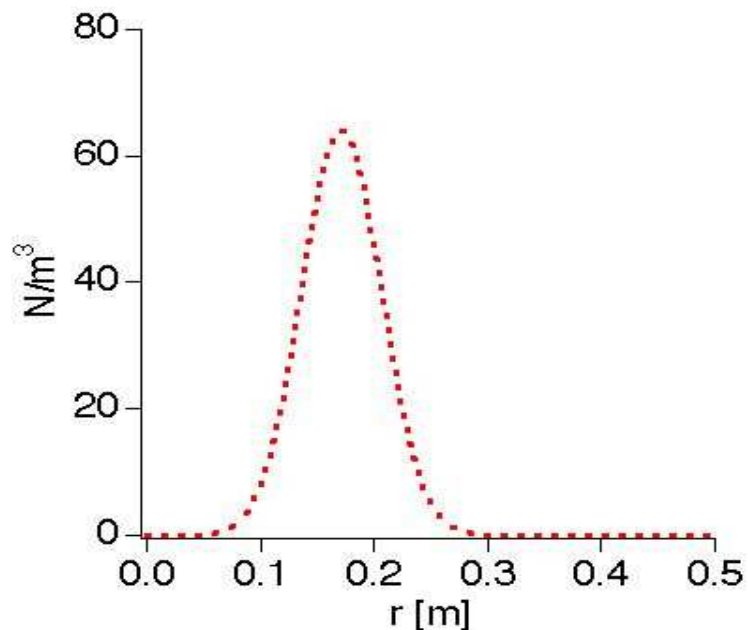
# Multiple helicity computations

# Three wave coupling generates localized electromagnetic torques .

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- $T_1 \approx \langle \tilde{J}_{k1} \times \tilde{B}_{k1} \rangle$  torque at the rational surface  $k1$ .
- If  $J_{k1} \approx B_{k2}B_{k3}$  is nonlinearly generated
- $T_1 \approx \langle B_{k2}B_{k3}B_{k1} \rangle$  **three wave coupling** [Hegna and Fitzpatrick]
- $T_1 + T_2 + T_3 = 0$
- The torque generated from nonlinear mode coupling is correlated with momentum transport.

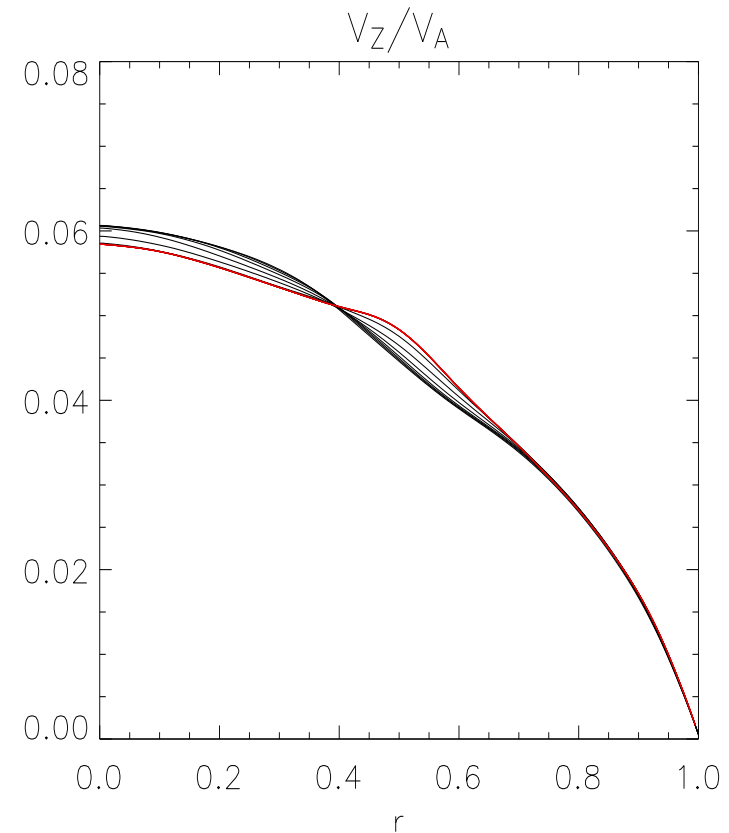
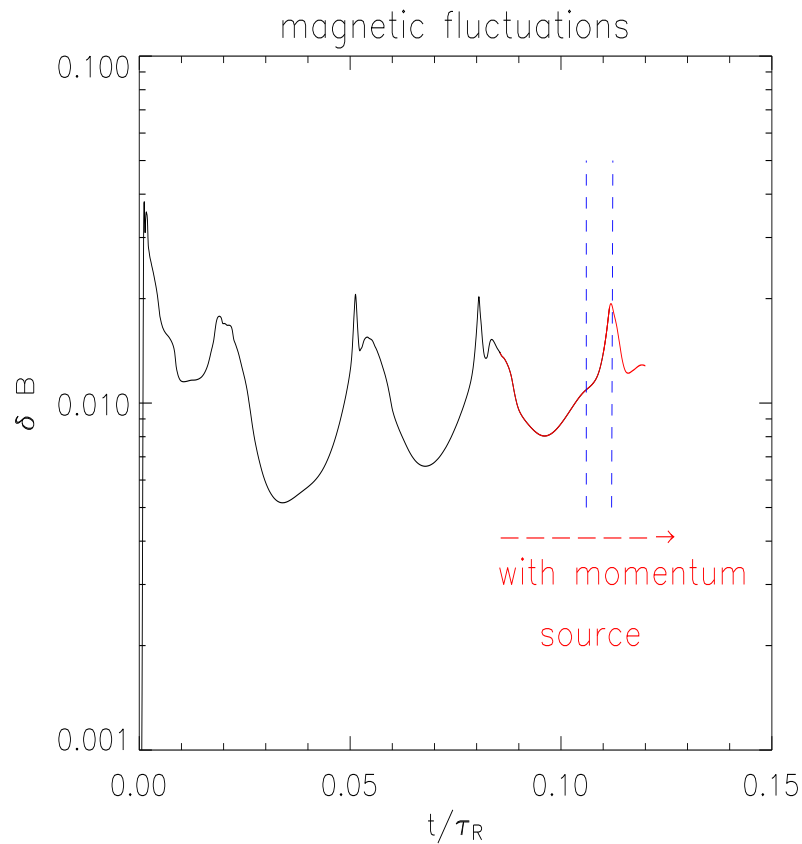
core mode  $\langle \tilde{J} \times \tilde{B} \rangle$  torque from experiment



# Flow profile is affected by the tearing fluctuations during sawtooth oscillations.

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$$S = 5 \times 10^4$$



- Higher Lundquist number computation show more pronounced changes of the flow profile during a sawtooth.

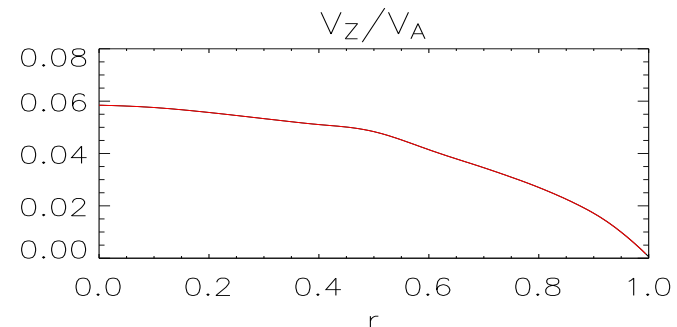
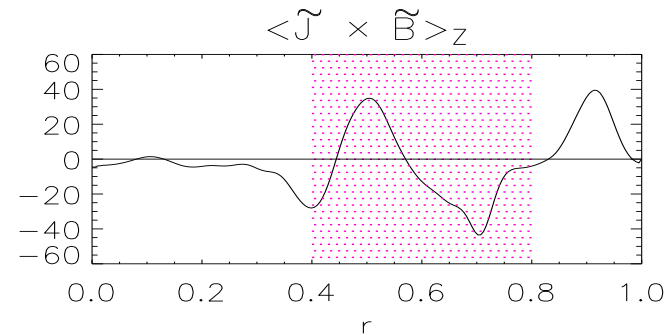
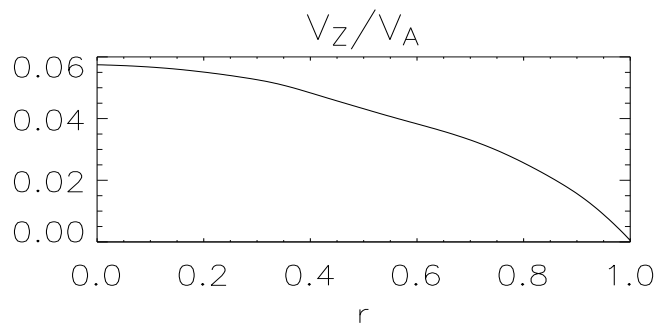
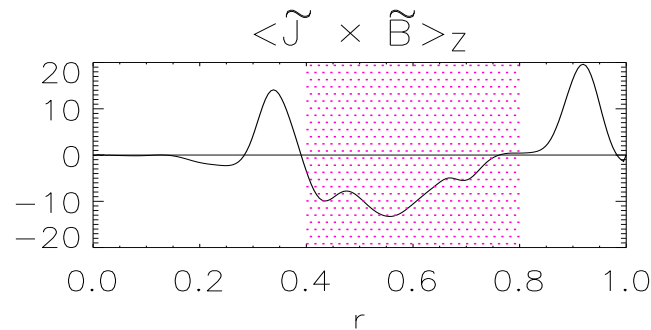
# There is correlation between the momentum transport and the fluctuation-induced forces.

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The fluctuation-induced forces at two times shown in previous slide.

low magnetic fluctuations

high magnetic fluctuations



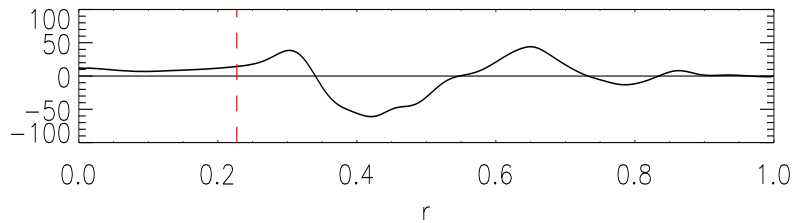
- The fluctuation-induced Lorentz force during the peak of the magnetic fluctuations results in flattening of the flow profile around  $r=0.4$ – $0.6$ .

# Total fluctuation-induced forces are the superposition of the single mode Lorentz forces.

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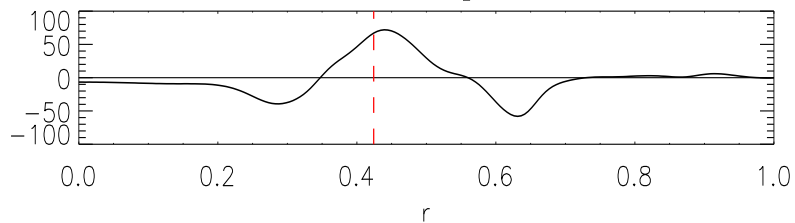
Axial  $\langle \tilde{J} \times \tilde{B} \rangle$  forces for core modes, (m,n)

$$\langle \tilde{J} \times \tilde{B} \rangle_z^{(1,-3)}$$



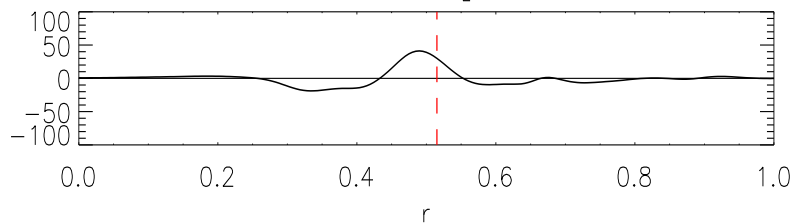
• (1, -3)

$$\langle \tilde{J} \times \tilde{B} \rangle_z^{(1,-4)}$$



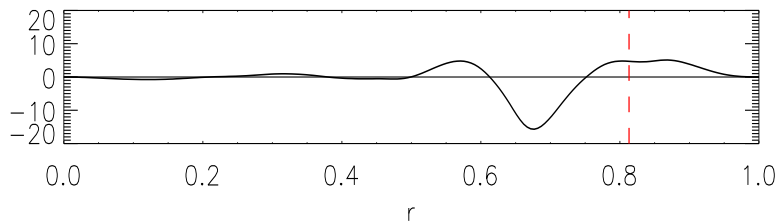
• (1, -4)

$$\langle \tilde{J} \times \tilde{B} \rangle_z^{(1,-5)}$$



• (1,-5)

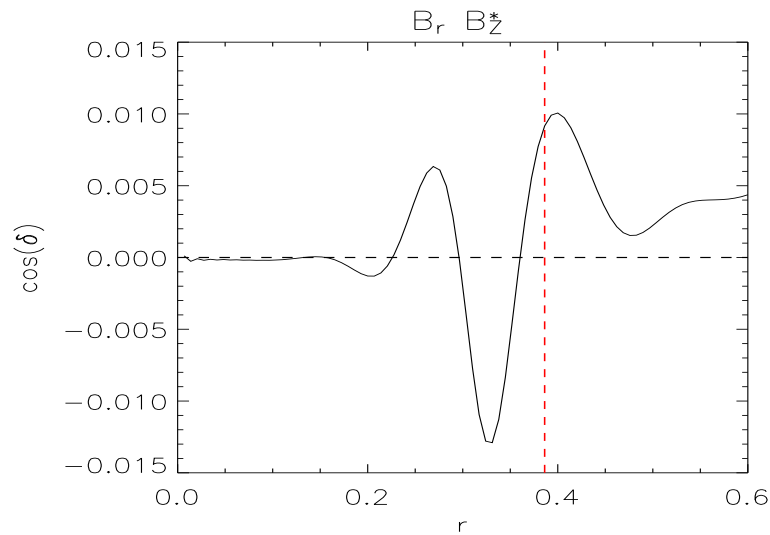
$$\langle \tilde{J} \times \tilde{B} \rangle_z^{(0,1)}$$



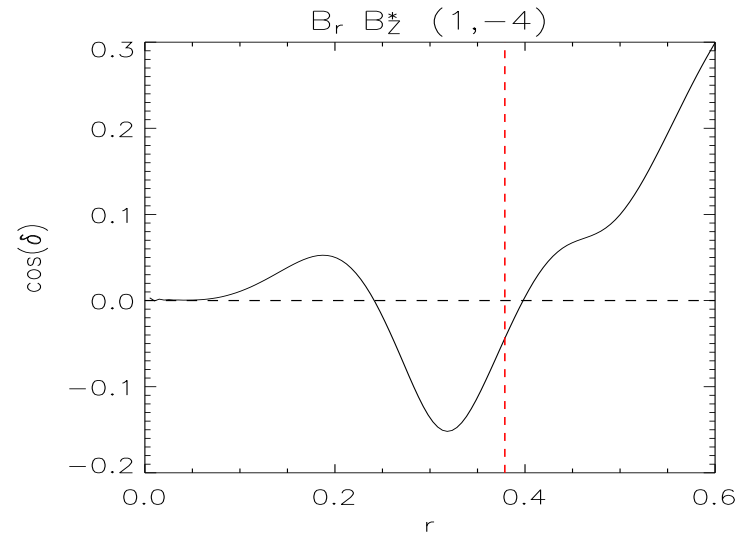
• (0, 1)

# Nonlinear mode coupling alters the radial phase dependency.

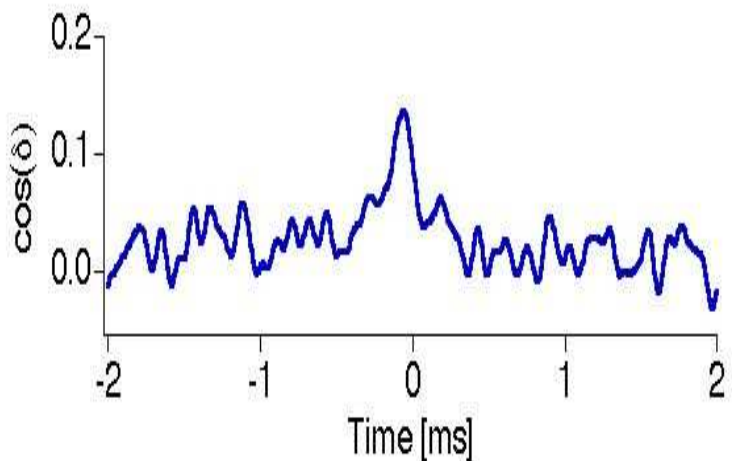
single helicity phase



multiple helicity phase



phase in experiment



- momentum transport is enhanced due to nonlinear mode coupling.

$$\langle \tilde{J} \times \tilde{B} \rangle = \frac{1}{r} \left[ r |\tilde{B}_z| |\tilde{B}_r| \cos(\delta) \right]'$$

# Conclusions

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- Quasilinear analytical calculations and single helicity computations show:  
momentum transport can occur due to single tearing mode  
(spontaneous reconnection)
- Nonlinear computations:  
momentum transport can arise from multiple mode coupling  
(coupled spontaneous and driven reconnection)
- Whether current-driven reconnection can transport momentum in astrophysical plasmas?