

# Reconnection and Line Tying

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# Why is this Important?

- Fieldlines in stellar & accretion disk coronae rooted in dense medium; what is effect on resistive stability?
  - energy storage
  - energy release
- Laboratory experiments have line tying: RWX, RSX
- Line tying has a dual role
  - Suppresses reconnection through magnetic tension
  - Forces reconnection through current sheet formation

# The Plan of This Talk

- Effect of line tied boundary conditions
- Reconnection in a line tied slab **with Yi-Min Huang**
- Example of current sheet formation

# Boundary Conditions

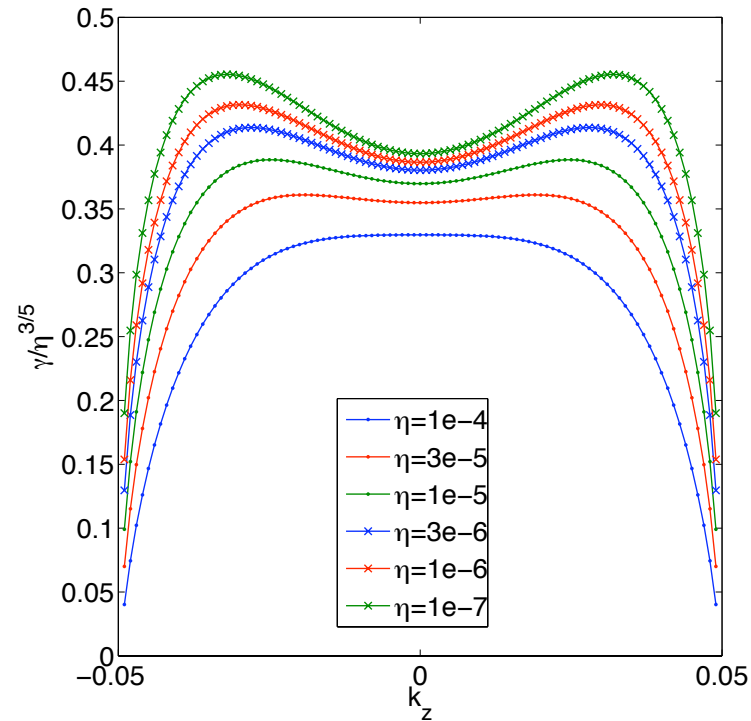
- Sheared field  $\mathbf{B} = \hat{y}B_y(x) + \hat{z}B_z(x)$
- Perturbations  $\delta Q = e^{\gamma t + i(k_y y + k_z z)} q(x)$
- Tearing modes localized to rational surfaces  $x_r$  where  $\mathbf{k} \cdot \mathbf{B} \equiv 0$
- Single Fourier component cannot satisfy line tied boundary conditions  $\delta \mathbf{V} = \delta B_z = 0$  at  $z = (0, L)$ .

# Recall Ideal Case

- Line tied kink stable if  $L < L_c \sim V_A/\gamma_p$
- If same here,  $L_c \propto \eta^{-3/5}$ .

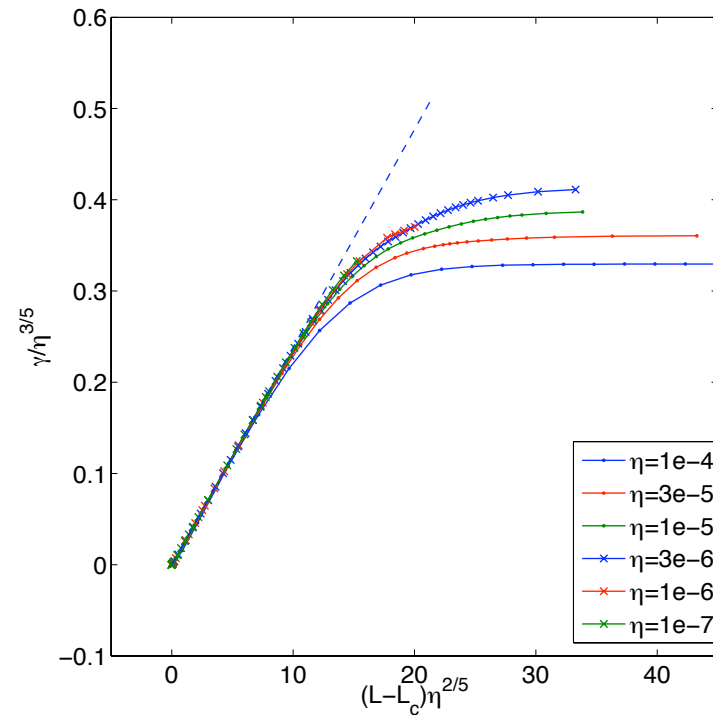
We will show that  $L_c \propto \eta^{-2/5}$ .

# Periodic System



Tearing mode growth rates  $\gamma_p$  for the Harris sheet equilibrium ( $B_y \propto \tanh x \ll B_z$ ) with periodic boundary conditions, for different  $\eta$ .

# Line Tied System



Scaled growth rates vs scaled length for Harris sheet with line tied b.c.. As  $L$  increases beyond critical length  $L_c$ ,  $\gamma \rightarrow$

$\gamma_p$ . At first  $\gamma \propto \eta$ , but as  $L$  increases there is a transition to

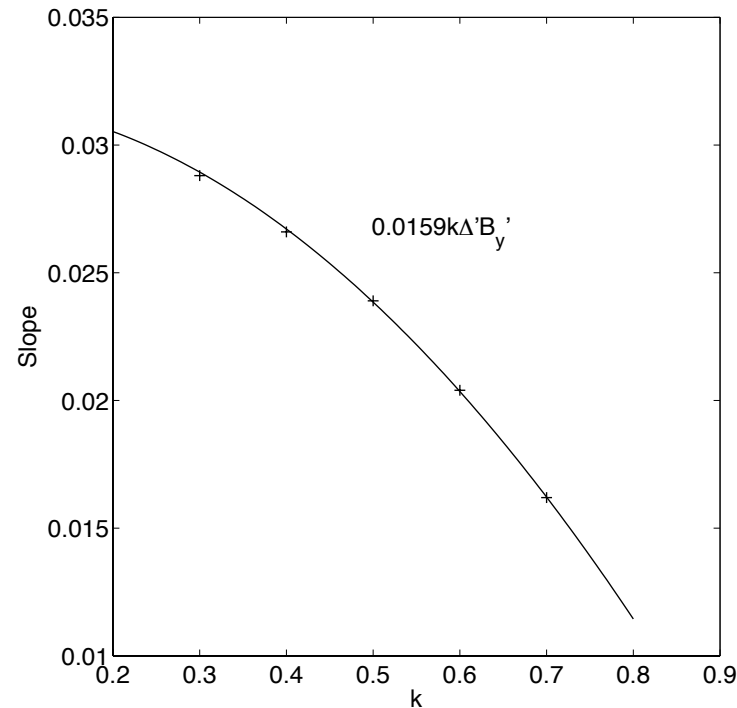
$\gamma \propto \eta^{3/5}$ . See also Delzanno & Finn.

# The Transition

- At long  $L$ , a pair of modes  $\pm k_z = \pm \frac{n\pi}{L}$  can combine to satisfy line tied B.C.
- Corresponding rational surfaces are  $x_n \sim \frac{n\pi}{k_y L} \frac{B_z}{B_y}$ .
- Tearing mode layer width  $\delta \propto (\Delta')^{1/5} \eta^{2/5}$ .
- $\gamma_p \propto (\Delta')^{4/5} \eta^{3/5}$

Transition is at  $L_t \propto \eta^{-2/5}$ , at which  $\delta \sim x_n$ . For  $L_c \ll L < L_t$ ,  
 $\gamma \propto \eta L$ .

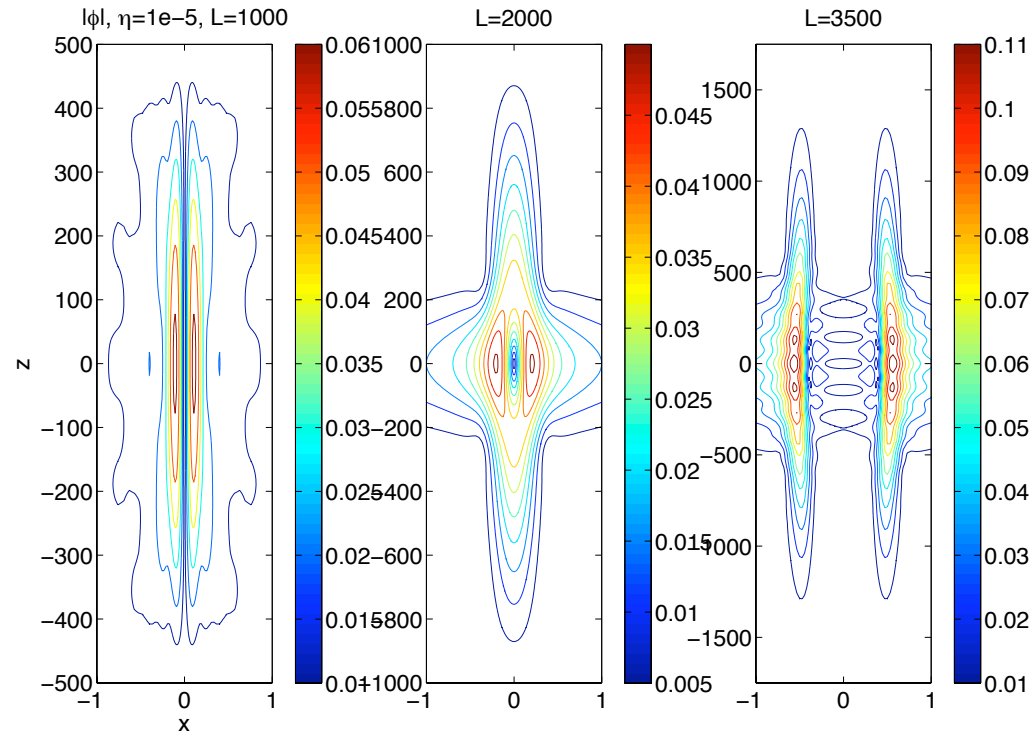
# The Slope



Find the slope of the  $\gamma - L$  relation such that  $\gamma = \gamma_p$  at  $L = L_t$ .

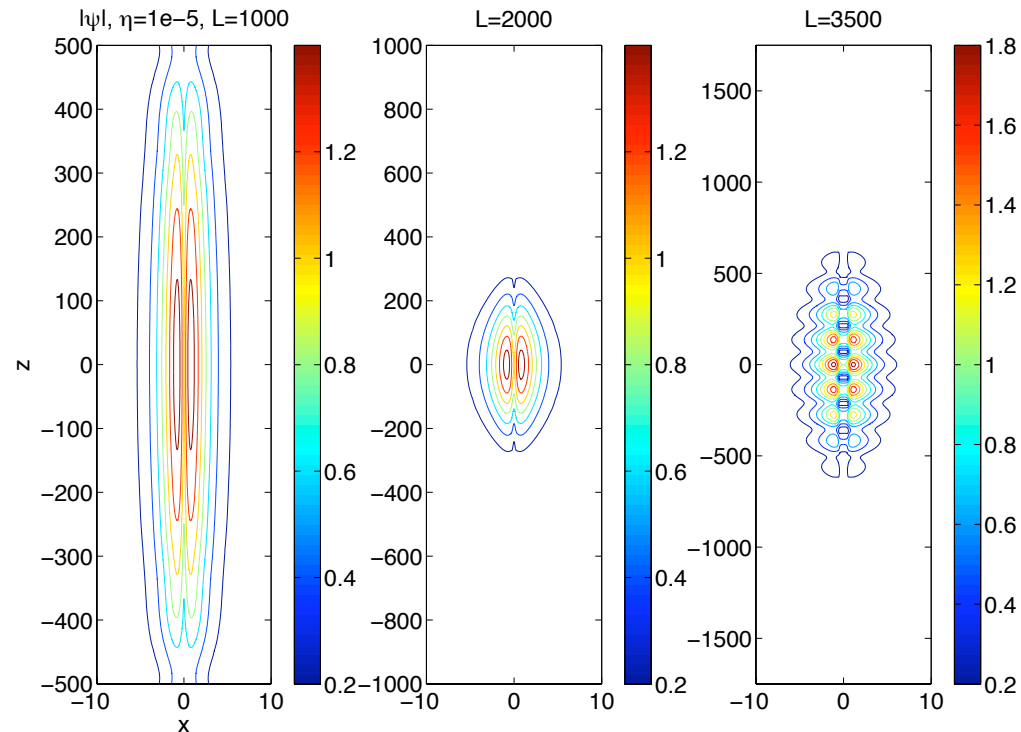
The fit is very good.

# The Stream Function



Stream function near  $L_c$ , near  $L_t$ , and  $L \gg L_t$ .

# The Flux Function



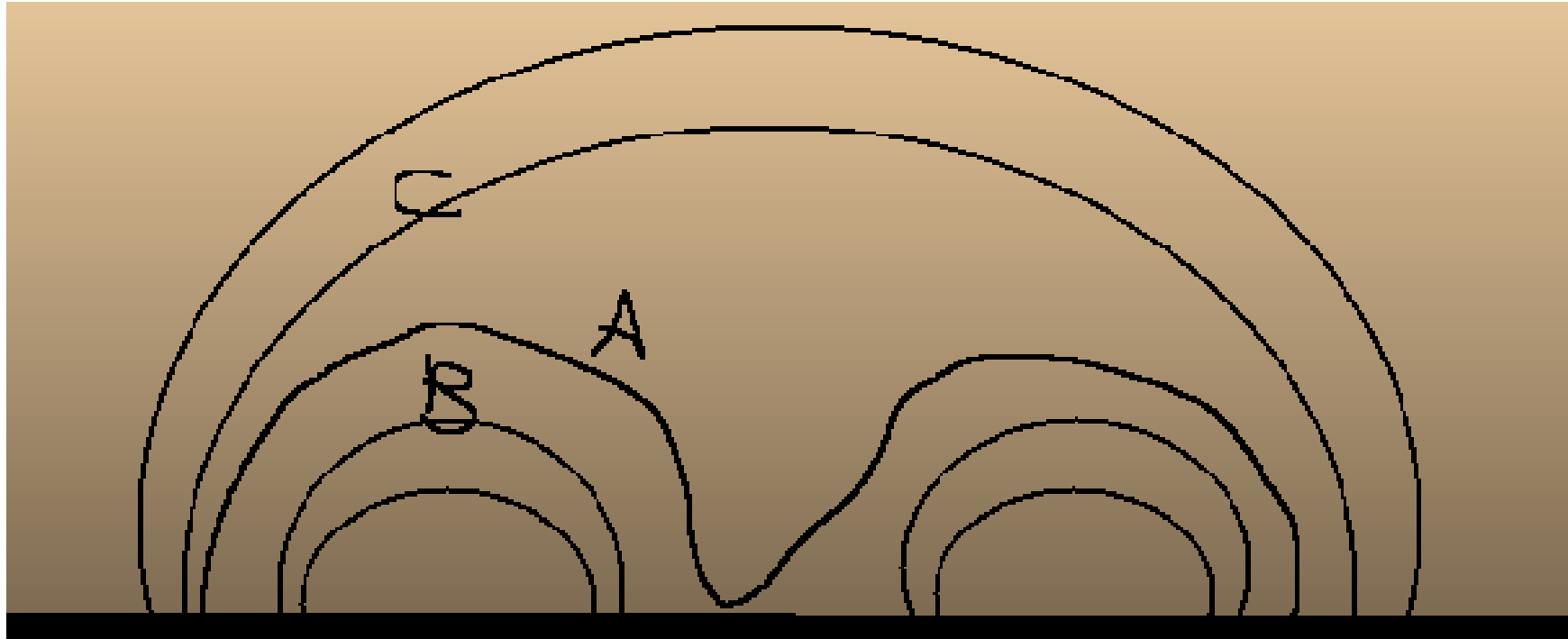
Flux function for the same 3 values of  $L$ . The mode approaches a constant length and high harmonics play an increasing role in localizing it.

# Is $\gamma \propto \eta$ Tearing?

- Growth rate depends on  $\Delta'$  and only  $k_y$  values for which the periodic system is unstable are unstable here.
- But equilibrium also evolves on resistive timescale.
- May be too slow to be interesting,
- Except for nonlinear evolution (high  $k$  structure).

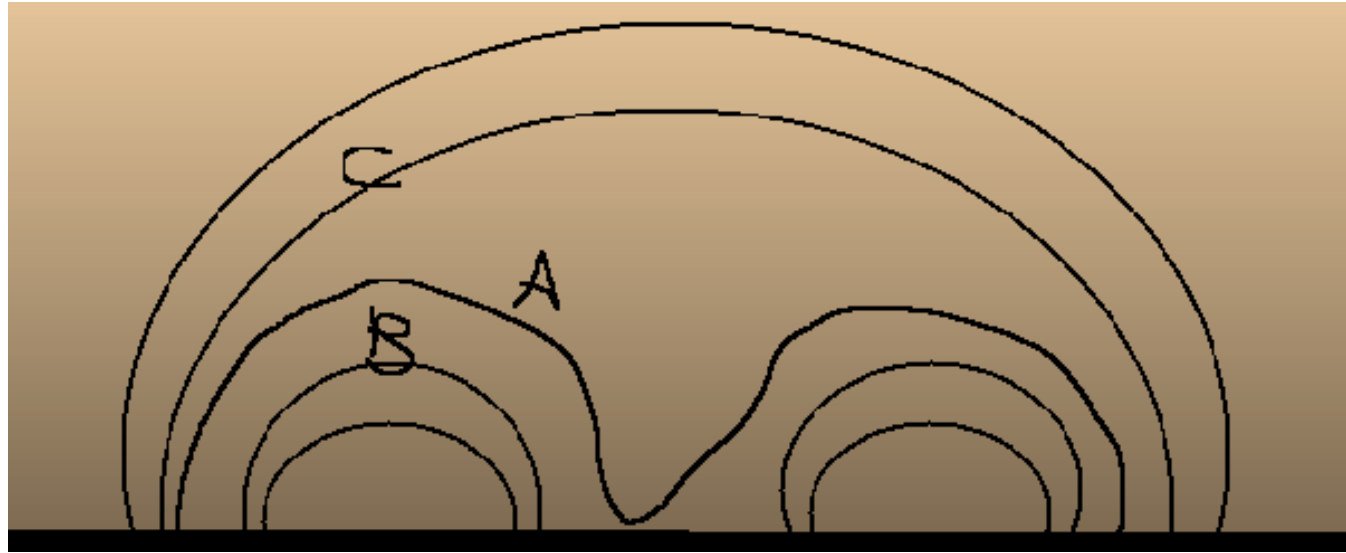
This strengthens the case that energy release in line tied systems begins with an ideal MHD process. Parker proposed current sheet formation is ubiquitous.

# Example



Line tied configuration with a separatrix (A). Line length doubles going from B type to C type across A.

# Current Sheet Formation



Shear the field by moving footpoints perpendicular to plane of screen & let field adjust to a new equilibrium. The shear per unit length will be  $\sim$  constant. Unless footpoint displacement doubles going from B to C, the shear will be discontinuous across A (see Low, also Zweibel & Proctor).

# Current Sheet Properties

- Viscously or resistively resolved:  $P_m \equiv \frac{\nu}{\eta} \sim \left( \beta_i \frac{m_i}{m_e} \right)^{1/2}$
- Equilibrium shear determines Poynting flux  $B_n \mathbf{v}_\perp \cdot \mathbf{B}_\perp$ .
  - What is the heating rate?
  - Is equilibrium shear ideal stable?

N.B. This current sheet has a simple topology but theorems on current sheet formation in a field extending between 2 plates do not apply. Separatrix is crucial & not unrealistic.

# Summary

- Linetying is an important astrophysical & lab boundary condition but precludes rational surfaces
- Resistive instability grows as  $S^{-1}$  for systems above a moderate length
- Rational surface behavior emerges in systems where tearing layers overlap for modes of periodic system
  - Transition length  $\propto S^{2/5}$ .
- Effective suppression of tearing strengthens case for MHD initiated energy release
- Current sheet formation across separatrices may be one way to achieve this.